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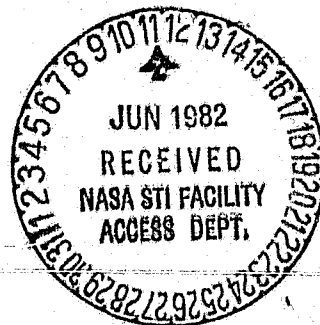
DEPARTMENT OF PHYSICS
SCHOOL OF SCIENCES AND HEALTH PROFESSIONS
OLD DOMINION UNIVERSITY
NORFOLK, VIRGINIA

Technical Report PTR-81-10

N₂ PRESSURE-BROADENED O₃ LINE WIDTHS
AND STRENGTHS NEAR 1129.4 CM⁻¹

By

G. E. Copeland, Principal Investigator
L. N. Majorana
C. N. Harward
and
R. J. Steinkamp



Progress Report
For the period October 25, 1980 - October 24, 1981

Prepared for the
National Aeronautics and Space Administration
Langley Research Center
Hampton, Virginia 23665

Under
Research Grant NAG1-1
James M. Hoell, Technical Monitor
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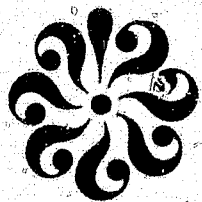
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N₂ PRESSURE-BROADENED O₃ LINE WIDTHS AND STRENGTHS NEAR 1129.4 CM⁻¹

By

G.E. Copeland¹, L.N. Majorana², C.N. Harward³,
and R.J. Steinkamp⁴

ABSTRACT

A Beer's Law experiment was performed with a tunable diode laser to find the N₂ pressure-broadening characteristics of a single O₃ absorption line at 1129.426 cm⁻¹ (v1 Band, J, Ka, Kc = 31, 1, 31 + 32, 0, 32) for N₂ pressures from 10 to 100 torr (O₃ pressure = 3.16 torr). SO₂ line positions were used for wavelength calibration. Line shapes were iteratively fitted to a Lorentz function. Results were $\delta(\text{HWHM in MHz}) = 47.44(\pm 5.34) \text{ MHz} + 1.730(\pm 0.088) \text{ MHz/torr} \cdot p(\text{torr})$ with $\rho = 0.9897$. This intercept compares well with the Doppler O₃ - O₃ broadened (at 3.16 torr) width of 44.52 Hz (ref. 1). Any difference is most likely in laser frequency uncertainty. This results in a HWHM line width of 0.044 cm⁻¹ atm⁻¹ at 760 torr and 285 K. The line strengths integrated over $\Delta\nu = 0.055 \text{ cm}^{-1}$ were found to be N₂ pressure dependent: $S(\text{cm}^{-2} \text{ torr}^{-1}) = 1.334 \times 10^{-4} p(\text{torr})^{-1.314}$ with $\rho = -0.9957$.

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INTRODUCTION

With the advent of semiconductor lasers and their recent application to spectroscopy (refs. 2-4), there has been a revolution in the study of infrared spectra. Use of the tunable diode laser (TDL) has permitted high-resolution ($<0.0001 \text{ cm}^{-1}$) studies of several molecules of high current interest (ClO , HNO_3 , SO_2 , CF_2Cl_2 , etc.). This study is concerned with the experimental determination of the pressure-broadened characteristics of a single O_3 line near 1129.4 cm^{-1} . This line has been suggested as a possible candidate for remote sensing experiments using either infrared heterodyne spectroscopy or differential lidar, since it is relatively well isolated from other atmospheric lines, has significant strength, and is relatively insensitive to temperature changes. A Beer's Law experiment is performed to determine the absorption coefficients and nitrogen (N_2) broadening parameters.

APPARATUS AND PROCEDURE

A simplified schematic of the apparatus is shown in figure 1. The radiation source is a PbSnTe TDL whose output is passed through collimating optics and a monochromator for mode selection. The output from the monochromator is divided by a beam splitter which directs 40 percent through a solid germanium etalon onto an LN_2 -cooled HgCdTe detector. The remaining 60 percent is split into 2 beams. One beam is directed through a reference gas cell to a second detector and the other through the 50-cm O_3 absorption cell to a third detector. A more detailed description of the apparatus is to be found in the report by Majorana (ref. 5).

This arrangement permits simultaneous recording of the O_3 absorption spectra, the etalon fringes for relative frequency calibration, and the SO_2 reference gas spectrum for conversion to absolute frequency (ref. 4). The detector output, the TDL current, and SO_2 and O_3 pressures were read by a 12-bit A/D and transferred to a 7-track tape by a PDP8/e minicomputer for later analysis. O_3 spectra were taken at 8 N_2 pressures (10-100 torr) at 285 K. SO_2 at 1 torr was used as the reference spectrum since a detailed atlas exists (ref. 6). The 1129.418 cm^{-1} line of SO_2 ($\nu_1, \nu_2, \nu_3, j, K_a, K_c$)

[1,0,0,33,3,31 + 0,0,0,34,2,32] was selected as the standard reference line because of its proximity to the O_3 line under study and the recent heterodyne confirmation of its position at $1129.41835 \text{ cm}^{-1}$.

THE DATA AND ITS ANALYSIS

The data were collected and processed in the same manner as described Majorana (ref. 5). Strip chart records were digitized and recorded on tape for later transfer to the ODU DEC System 10 computer. After input of data into the proper data format, each of the N_2 -broadened O_3 profiles was processed using the program QUICK. One of the output products from program QUICK is a plot of the absorption coefficient as a function of wavenumber, as well as generation of data files containing the coefficient and the transmittance functions.

For the foreign gas broadening study done in this research, the proper broadening function is a Lorentz function. Therefore, the Voigt function-fitting routine developed by Majorana and Copeland (ref. 5) cannot be used efficiently. A new program, BROAD (listed in the Appendix) was developed to fit in a nonlinear least squares sense a Lorentzian to the absorption coefficient: that is, a fit of the data is made to the function

$$y = A_4 + \frac{A_1}{(\omega - A_2)^2 + \left(\frac{A_3}{2}\right)^2} \quad (1)$$

where A_1 is the numerator of the Lorentzian, A_2 is the center frequency (in cm^{-1}), A_3 is the half-width at half maximum (HWHM) of the profile, and A_4 is the background.

The program BROAD makes extensive use of the data-fitting routines developed by Bevington (ref. 8) and provides a file for plotting its fitted function together with the data for the user. The program begins by asking the user the name of the input data file. The program reports the number of

data points, the date the data was taken, the total ($O_3 + N_2$) pressure in torr, and the minimum and maximum values of the absorption coefficient. The subroutine CHOP is called, and the user may reduce the size on high and low wavenumber ends of the spectrum to analyze. The total wavenumber interval across the data is reported, and the user may normalize the data if needed. The function AREA is called, and the integral of the absorption coefficient over wavenumber is calculated, i.e., the strength in units of $cm^{-2}torr^{-1}$. The user may select any of three different procedures in which to weight the data during the fitting procedure (MODE = +1,0,-1) which correspond to the Bevington weight convention (ref. 8).

The program BROAD was designed to fit Lorentzians to either convex or concave data sets, i.e., A_4 near zero or A_4 maximum. The program automatically detects each case and, assuming only one spectral line, then makes initial guesses at the parameters A_1 , A_2 , A_3 , A_4 and at their initial movements DELTAA(1-4). These are reported to the user. The program uses the grid-search method, subroutine GRIDLS, and the DO loop 100 is entered. After each exit from GRIDLS, the reduced chi square statistic is compared with the previous value. If it has not changed by more than two percent, iteration is stopped and the program reports the final chi square, the number of iterations, the final values of the parameters A_1 to A_4 , and estimates of their errors. Finally, the total pressure in torr and the half-width at half maximum in both cm^{-1} and MHz are reported. The user may then select to see a recap of the data and the best fit and plot such as a graphics terminal. An example output of program BROAD is shown in the appendix.

RESULTS

These results are the first report to date of ozone line width broadening in the ν_1 band using N_2 as the foreign gas. Table 1 lists the N_2 pressure, the fitted half-widths, and the line strengths for the case where the initial ozone pressure was 3.16 torr. Figures 2 through 10 show the Lorentz fit to the absorption coefficient from the program BROAD.

The half-widths as a function of pressure are shown in figure 11. We find they can be represented by

$$\delta(\text{HWHM in MHz}) = A + B p_{N_2} \quad (2)$$

with

$$A = 47.4 \pm 6.3 \text{ MHz}$$

and

$$B = 1.730 \pm 0.01 \text{ MHz/torr}$$

with a linear correlation coefficient of 0.9897. The intercept value 47.4 ± 5.3 MHz corresponds to the $O_3 - O_3$ broadened width at 3.16 torr O_3 . This self-broadening half-width compares well with the results of Majorana et al. (ref. 1), who found a width of 44.52 MHz. The two values agree well within error bounds. Any significant difference could be traced to frequency instability in the diode laser.

The strength of the absorption line is defined as the integral over wavenumber of the absorption coefficient; that is, strength S is given by

$$S(\text{cm}^{-2}\text{torr}^{-1}) = \int k(\nu) d\nu \quad (3)$$

Generally, S is a constant for self-broadened data; however, we find S to be N_2 pressure dependent. Figure 12 shows the strength for this transition as a function of N_2 pressure. We find that the strength can be represented as a power function of the N_2 pressure, i.e., $S = A p_{N_2}^B$, where $A = 1.334 \times 10^{-4}$ and $B = 1.314 \pm 0.049$ with a correlation coefficient of -0.9957. Since -1.34 is essentially -4/3, then we find the strength varies as pressure of nitrogen raised to the -4/3 power. This curious result may be due to several effects: chemical decomposition of O_3 due to collisions with N_2 or formation of weak Van der Waals molecules with N_2 , i.e., $O_3 + N_2 \rightarrow O_3N_2$ and the resulting modification of residual ozone concentration.

APPENDIX

PROGRAM BROAD

This program is designed to perform the main data processing tasks as described in detail in the section "The Data and Its Analysis." This program is written in Fortran-10 and runs on a Dec System-10 time shared computer under the Tops10 operating system. Extensive use is made of the following data handling routines selected from those of Bevington (ref. 3) and are not listed here:

```
GRIDLS(X,Y,SIGMAY,NPTS,NTERMS,MODE,A,DELTA,SIGMAA,YFIT,  
CHISQR)
```

```
INTEG(X,Y,NPTS,1,X1,X2,SUM)
```

```
MATINV(ARRAY ,NTERMS,DET)
```

```
FUNCTION FCHISQ(Y,SIGMAY,NPTS,NFREE,MODE,YFIT)
```

```
FUNCTION AREA(X,Y,NPTS,NTERMS)
```

LISTING

```
PROGRAM BROAD
```

```
WRITTEN BY
```

```
DR. G. E. COPELAND
```

```
DEPARTMENT OF PHYSICS AND GEOPHYSICAL SCIENCES
```

```
OLD DOMINION UNIVERSITY
```

```
NASA HETERODYNE GRANT
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```
DIMENSION YLABEL(16),XLABEL(10),TITLE(35),SIGMAX(400),  
1SIGMAY(400)
```

```
DIMENSION X(400),Y(400),YFIT(400),A(4),SIGMAA(4)
```

```
DIMENSION DELTAA(4)
```

```
DOUBLE PRECISION FILE
```

```
SPEED OF LIGHT IN CM PER SECOND
```

```
C=2.9979250E10
```

```
DATA FOR THE LABELS ON THE GRAPH
```

```
DATA XLABEL/'W','A','V','E','N','U','M','B','E','R'/
```

```

DATA YLABEL/'A','B','S',',',
1'C','O','E','F','F',',',',B','A','S','E',',',',E'/
DATA TITLE/'O','3',',',',L','I','N','E',',',',A','T','1','1',
1'2','9',',',',4',4',3',C',M',-',1',',',L','O','R','E','N',
2,'T','Z',',',',F','I','T'/

```

```

C
DATA YLAB3/'R','E','S','I','D','U','A','L','S'/
TYPE 92
92 FORMAT(' L O R E N T Z I A N FITTING FUNCTION')
WRITE(5,10010)
ACCEPT 10020 ,FILE
OPEN(UNIT =21,DEVICE='DSK',ACCESS='SEQIN',FILE=FILE)
READ(21,*)NPTS,K2,K3,K4,P3
IF(NPTS.GT.400) TYPE 41
41 FORMAT(' TOO MANY DATA POINTS ENLARGE THE DIMENSIONS')
IF(NPTS.GT. 400) STOP
TYPE 222,NPTS
222 FORMAT(' NUMBER OF DATA PAIRS =',I4)
TYPE 451,K2,K3,K4
451 FORMAT(1X,' MONTH =',I2,' DAY =',I3,' YEAR =',I4)
C
TYPE 4519,P3
4519 FORMAT(' PRESSURE IS ',F7.2,' TORR')
C
C READ THE DATA IN THE FILE INTO X AND Y ARRAYS
C
CALL READDT(X,Y,NPTS)
C
C
DO 31 I=1,NPTS
C TYPE 32 ,I,X(I),Y(I)
32 FORMAT( 1X,I3,1X,F8.3,1X,F8.3)
31 CONTINUE
C
C
CLOSE(UNIT=21,DEVICE='DSK',ACCESS='SEQIN',FILE=FILE)
C
C CALL ROUTINE TO CHOP OUT 1ST AND LAST OF DATA TO GET RID OF OUTER LINES
C
CALL CHOP(NPTS,X,Y)
C
C
C ASK IF DATA IS TO NORMALIZED
C
TYPE 77
77 FORMAT(' DO YOU WANT THE DATA NORMALIZED?<Y=1,N=0>'$)
ACCEPT *,ANNOR
IF ( ANNOR .EQ. 0) GOTO 79

```

```

C
C      YES NORMALIZE IT
C
C      ANSW=AREA(X,Y,NPTS,3)
C
C
C      SCALE THE Y ARRAY NOW
C
C      DO 732 I=1,NPTS
732    Y(I)=Y(I)/ANSW
C
C
C      TYPE 765,ANSW
765    FORMAT(' INTEGRAL OVER THE DATA IS ',1PE)
79    OPEN(UNIT=22,DEVICE='DSK',ACCESS='SEQOUT',FILE='PLTTEK.DAT')
      NSHET=1
      NGRP=2
      WRITE(22,10190) NSHET
      INPUT MODE +1,0,-1.
C
C      MODE=+1 IS THE INSTRUMENTAL
C      MODE = -1 IS STATISTICAL
C      MODE =-1
C      MODE =0 NO WEIGHTING
C
40    WRITE(5,10100)
      MODE=-1
      ACCEPT *,MODE
C      CALCULATE MEAN IN XAND Y-
      YMEAN=0.0
      XMEAN=0.0
      DO 50 I=1,NPTS
      SIGMAX(I)=0.0
      XMEAN=X(I)+XMEAN
      SIGMAY(I)=0.0
      YMEAN=YMEAN+Y(I)
50    CONTINUE
      XMEAN=XMEAN/FLOAT(NPTS)
      YMEAN=YMEAN/FLOAT(NPTS)
C      CALCULATE 'SD' OF 'Y' AND 'SD' OF 'X' (SDY _SDX)
      SDX=0.0
      SDY=0.0
      DO 70 I=1,NPTS
      SDX=(X(I)-XMEAN)**2+SDX
      SDY=(Y(I)-YMEAN)**2+SDY
      SIGMAX(I) = SQRT( (X(I)-XMEAN)**2 )
      SIGMAY(I) = SQRT((Y(I)-YMEAN)**2)
70    CONTINUE
      SDX=SQRT(SDX/FLOAT(NPTS-1))
      SDY=SQRT(SDY/FLOAT(NPTS-1))

```

```

C      WE HAVE SDY SDX TO PROVIDE START UP VALUES FOR
C      THE GRID SEARCH
C
      YMAX=Y(1)
      DO 80 I=2,NPTS
      IF( YMAX .LT. Y(I) ) YMAX=Y(I)
80     CONTINUE
      NTERMS=4
      YMIN=Y(1)
      DO 81 I=2,NPTS
      IF( YMIN .GT. Y(I) ) YMIN=Y(I)
81     CONTINUE
C
C
C      G U E S S   R O U T I N E
C
      A(4)=( Y(1) +Y(NPTS) )*.5
      TYPE *,YMIN,YMEAN,YMAX,XMIN,XMEAN,XMAX
      A(2)=XMEAN
C
      IF(Y(1) .LT. YMEAN ) GOTO 811
C
C      CURVE IS CONCAVE
      ICON=1
      TYPE 877
877     FORMAT('          CURVE IS CONCAVE')
      YMID=Y(1)-( Y(1)-YMIN)*0.5
      DO 992 I=1,NPTS
      IF( Y(I) .LT. YMID) XL=X(I)
      IF( Y(I) .LT. YMID) ILEFT=I
      IF( Y(I) .LT. YMID) GOTO 993
992     CONTINUE
      STOP '992'
993     DO 994 I=ILEFT,NPTS
      IF( Y(I).GT. YMID) XR=X(I)
      IF( Y(I) .GT. YMID) GOTO 995
994     CONTINUE
      STOP '994'
811     CONTINUE
C
      CURVE IS CONVEX
      TYPE 878
878     FORMAT(' CURVE IS CONVEX')
C
      YMID=Y(1)+(YMAX-Y(1))*0.5
      DO 8121 I=1,NPTS
      IF( Y(I) .GT. YMID) XL=X(I)
      IF( Y(I) .GT. YMID) ILEFT=I
      IF( Y(I) .GT. YMID) GOTO 8131
8121    CONTINUE
      STOP '8121'

```

```

8131 DO 814 I=ILEFT,NPTS
      IF( Y(I) .LT. YMID) XR=X(I)
      IF( Y(I) .LT. YMID) GOTO 995
814  CONTINUE
      STOP '814'
995  A(3)=XR-XL
C
C      A3 IS FWHM
C
      A(1)=(0.25* A(3)**2+(XMEAN-A(2))**2)*( YMAX-A(4) +0.5)
      IF( Y(1).GT. YMIN) A(1)=-A(1)
C
C      GOT A1,A2,A3,A4
C
C      TYPE *,A
C
C      NOW HAVE FINISHED THE
C      GENERATION OF GUESSES ON A1,A2,A3,A4 FOR A LORENTZIAN FUNCTION
C       $Y=A_4 + A_1 / ( (X-A_2)^2 + (A_3/2)^2 )$ 
C
      WRITE(5,10120)
      WRITE(5,10110)(I,A(I),I=1,NTERMS)
C      FINISHED STARTING VALUES
C      NOW FIND THE INCREMENTS ON A(1),A(2),A(3)
C      SET STEP SIZE IN MEAN = 6SDX/20
      DELTAA(3)=A(3)/10.
C      SET STEP SIZE ON COEFFICIENT A1
      DELTAA(1)=A(1)/10.
C      SET STEP SIZE ON COEFFICIENT A2
      DELTAA(2)=(XMEAN+SDX)/10.0
C      GIVE GUESS ON DELTAA(4)-----BACKGROUND
      DELTAA(4)=A(4)/10.0
C
C      ENTER LOOP
C      ITERATE 20 TIMES AT MOST
      KNT=0
      DO 100 K=1,15
      CALL GRIDLS(X,Y,SIGMAY,NPTS,NTERMS,MODE,A,DELTAA,
1      SIGMAA,YFIT,CHISQR)
      TYPE 887,K,CHISQR
      IF(K .GT. 1) GOTO 90
      SAVE=CHISQR
      SAVIT=SAVE
      KNT=1
      GOTO 100
90    XCHI=100.0*(CHISQR-SAVE)/CHISQR
      IF( XCHI .GT. 0.0 ) GOTO 110
C
CRITERION OF CONVERGENCE IS CHISQR DOES NOT CHANGE BY 1 PER CENT

```

```

C
C      OR IF CHISQR IS LESS THAN 0.02 OF THE 1ST CHISQR FOUND
C
C      BEVINGTON PAGE 212
C
      IF( ABS(XCHI) .LT. 0.5 ) GOTO 110
      IF( CHISQR .LT. 0.020*SAVIT) GOTO 110
      SAVE = CHISQR
      KNT=KNT+1
100    CONTINUE
110    KNT=KNT+1
      WRITE(5,10130) KNT
C      RETURNED VALUES ARE A,SIGMAA,YFIT,CHISQR
      WRITE(5,10140)
      DO 120 I=1,NTERMS
      WRITE(5,10150),I,A(I),SIGMAA(I)
120    CONTINUE
      WRITE(5,10160),CHISQR
      WRITE(5,10161),P3
10161  FORMAT('    TOTAL PRESSURE IS =',F6.2,' TORR')
      A3K=A(3)*.5
      A3MHZ=A3K*C/1.E+6
      TYPE 345 ,A3K,A3MHZ
345    FORMAT(1X,'HWHM =',F9.6,' CM-1 OR ',F12.4' MHZ')
130    WRITE(5,10170)
C      ANS1=0
      ACCEPT *,ANS1
      IF(ANS1.EQ.1)GO TO 150
140    WRITE(5,10180)
      ACCEPT *,ANS2
      IF(ANS2.EQ.0) STOP
      WRITE(22,10190) NGRP
      WRITE(22,10190),NPTS
      WRITE(22,10080)(YLABEL(I),I=1,16)
      WRITE(22,10080)(XLABEL(I),I=1,10)
      WRITE(22,10200)(TITLE(I),I=1,35)
      WRITE(22,3451),A3K,A3MHZ,P3
3451  FORMAT(1X,'HWHM ',F9.6,' CM-1 ',F7.2' MHZ P= 'F6.2' TORR')
C      OUTPUT TO TEKTRONIX PLOTTING FILE
C
      DO 812 I=1,NPTS
      WRITE(22,10220) X(I),Y(I)
812    CONTINUE
C
C      WRITE FITTED DATA AND GENERATE NEW X VALUES
C
C
C      WRITE OUT THESE X AND YFIT
      DO 813 I=1,NPTS
813    WRITE(22,10220) X(I),YFIT(I)
99    CONTINUE

```



```

CLOSE(UNIT=22,DEVICE='DSK',ACCESS='SEQOUT',FILE='PLTTEK.DAT')
STOP
C      OUTPUT A DATA MATRIX IN TERMINAL
150    CONTINUE
      Q=0.0
C
C      FIND AVG OF ABS(% ERROR)
C
      WRITE(5,10240)
      DO 88 I=1,NPTS,10
      Z=( (Y(I)-YFIT(I))/YFIT(I) ) * 100.0
      Q=Q+ABS(Z)
88     WRITE(5,10250) X(I),Y(I),YFIT(I),Z
      Q=10.*Q/FLOAT(NPTS)
      WRITE(5,10230)Q
      WRITE(5,10240)
      GO TO 140
887    FORMAT(' FINISHED ITERATION ',I2,' CHI SQ =',1PE)
10010  FORMAT(1X,'WHAT IS INPUT FILE NAME<CONTAINS OUTPUT FROM
QUICK>?')
10020  FORMAT(A10)
10060  FORMAT(I)
10080  FORMAT(40A1)
10090  FORMAT(1X,40A1)
10100  FORMAT(1X,'WHICH MODE DO YOU WISH TO FIT DATA',/,
1      1X,'+1=INSTRUMENTAL FITS= 1/VARIANCE',/,
2      1X,'0= NO WGHT',/,
3      1X,'-1= STATISTICAL=1/Y = ',/,1X,'?'$)
10110  FORMAT(5X,' A(',I1,' ) =',1PE)
10120  FORMAT(10X,' S T A R T U P V A L U E S')
10130  FORMAT(1X,'THERE WERE ',I3,' ITERATIONS')
10140  FORMAT(1X,'USING Y=A4+A1/( (X-A2)**2 + 0.25*A3 **2)
1',/,1X, 'YOU OBTAIN:')
10150  FORMAT(5X,'A(',I1,' ) =',1PE,'+/- ',1PE)
10160  FORMAT(5X,'CHI SQUARED = ',1PE)
10170  FORMAT(1X,/,1X,'DO YOU WANT A DATA RECAP? YES=1,NO=0?'$)
10180  FORMAT(2X,'DO YOU WANT TO PLOT DATA ON GRAPHICS ?(YES=1,NO=0)'$)
10190  FORMAT(2X,I)
10200  FORMAT(40A1)
10210  FORMAT(5X,'A1-4=',4(1X,1PE9.2) )
10220  FORMAT(2X,1PE,2X,1PE)
10230  FORMAT(1X,'MEAN OF THE ABSOLUTE VALUE OF THE % ERROR =',F)
10240  FORMAT(2X,/,/,8X,'X-DATA',8X,'Y-DATA',8X,
1'YFIT',8X,'PER CENT DIFF')
10250  FORMAT(1X,3(1PE,3X),OPF10.2)
      END
C
C-----
C      FUNCTION FUNCTN(X,I,A)
C
C

```

C TAKEN FROM
 C DATA REDUCTION AND ERROR ANALYSIS
 C FOR THE
 C PHYSICAL SCIENCES
 C BY
 C PHILIP R. BEVINGTON
 C MCGRAW-HILL BOOK CO 1969
 C PAGE 214

C G E COPELAND JUNE 1980
 C DEPARTMENT OF PHYSICS AND GEOPHYSICAL SCIENCES
 C OLD DOMINION UNIVERSITY
 C NORFOLK, VA. 23508

C PURE LORENTZIAN FUNCTION

C $Y=A_4+(A_1)/((X-A_2)**2+(A_3/2)**2)$

C PLUS A BACKGROUND

C DIMENSION X(1),A(1)

10

 XI=X(I)

30

 Z=XI-A(2)

 Z2=Z*Z

 FUNCTN=A(1)/(Z2+(A(3)*A(3)*0.25))

 FUNCTN=FUNCTN+A(4)

50

 RETURN

 END

C

C

C-----

C

 SUBROUTINE READDT(X,Y,N)

 DIMENSION X(1),Y(1)

C10

 READ(21,211)KK,N,K2,K3,K4,K10,TAPENO,P3

C211

 FORMAT(1X,I2,1X,I4,1X,3(I2,1X),I1,1X,A5,1X,F)

 READ(21,END=22,*)(X(I),Y(I),I=J,J+4),J=1,N,5)

334

 FORMAT(10(1X,1PE8.2))

C

 WRITE(5,334)((X(I),Y(I),I=J,J+4),J=1,N,5)

 RETURN

22

 CONTINUE

 TYPE 33

33

 FORMAT('RAN OUT OF DATA IN READDT---RECAP FOLLOWS')

 TYPE 34 ,N

34

 FORMAT('NUMBER OF DATA POINTS =',I)

 WRITE(5,*)(X(I),Y(I),I=J,J+4),J=1,N,5)

 RETURN

```

      END
C
C-----
C
      SUBROUTINE CHOP(N,X,Y)
      DIMENSION X(1),Y(1)
      DIMENSION XSAV(400),YSAV(400)
C
C      CHOP OUT PARTS OF THE ARRAY AND REPLACE IT IN THE ORIGINAL ARRAYS
C
      NSAV=N
      DO 10 I=1,NSAV
      XSAV(I)=X(I)
10     YSAV(I)=Y(I)
      YMIN=Y(1)

      DO 20 I=2,NSAV
      IF(YMIN .GT. Y(I)) IMIN=I
      IF(YMIN .GT. Y(I)) YMIN=Y(I)
20     CONTINUE
      TYPE 100 ,YMIN,IMIN
100    FORMAT(' MINIMUM VALUE OF Y ='F,' AT POINT ',I3)
      YMAX=Y(1)
      DO 201 I=2,NSAV
      IF( Y(I) .GT. YMAX) IMAX=I
      IF( Y(I) .GT. YMAX) YMAX=Y(I)
201    CONTINUE
      TYPE 2011,YMAX,IMAX
2011   FORMAT(' MAXIMUM VALUE OF Y ='F,' AT POINT ',I3)
109    TYPE 110
110    FORMAT(' DO YOU WANT TO CHOP PARTS OF THE ARRAY OFF?<Y=1,NO=0>')
      ACCEPT *,ANS
      IF(ANS.EQ.0) RETURN
      IF(ANS.EQ.1) GOTO 150
      GOTO 109
150    CONTINUE
      TYPE 160
160    FORMAT(' INPUT   OF LEFT DATA POINT TO START NEW ARRAY?')
      ACCEPT *,IST
      TYPE 161
161    FORMAT(' INPUT   OF DATA POINT ON RIGHT TO STOP THE ARRAY?')
      ACCEPT *,ISP
      IF( ISP.LT.IST) GOTO 109
C
C      REPORT THE DIFFERENCE IN FREQUENCY BETWEEN THE 2 LIMITS
C
      DIFF=X(ISP)-X(IST)
      TYPE 1601,DIFF
1601   FORMAT(' WAVENUMBER CHANGE ACROSS THIS SAMPLE ='F8.7,' CM-1')
      DO 30 I=IST,ISP
      X(I-IST+1)=XSAV(I)

```

```

Y(I-IST+1)=YSAV(I)
N=ISP-IST
RETURN
END

```

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RUN BROAD

L O R E N T Z I A N FITTING FUNCTION

WHAT IS INPUT FILE NAME<CONTAINS OUTPUT FROM EARL>?
FIL5

NUMBER OF DATA PAIRS = 200

MONTH =12 DAY = 19 YEAR = 79

PRESSURE IS 53.59 TORR

MINIMUM VALUE OF Y = -0.0000024 AT POINT 180

MAXIMUM VALUE OF Y = 0.0000723 AT POINT 93

DO YOU WANT TO CHOP PARTS OF THE ARRAY OFF?<Y=1,N=0>

1

INPUT # OF LEFT DATA POINT TO START NEW ARRAY?

20

INPUT # OF DATA POINT ON RIGHT TO STOP THE ARRAY?

180

WAVENUMBER CHANGE ACROSS THIS SAMPLE =.0484009 CM-1

DO YOU WANT THE DATA NORMALIZED?<Y=1,N=0>

1

INTEGRAL OVER THE DATA IS = 8.1862793E-07

WHICH MODE DO YOU WISH TO FIT DATA

+1=INSTRUMENTAL FITS= 1/VARIANCE

0= NO WGT

-1= STATISTICAL=1/Y =

?

0

CURVE IS CONVEX

S T A R T U P V A L U E S

A(1) = -1.5973438E-03

A(2) = 1.1294275E+03

A(3) = 8.4991455E-03

A(4) = 3.6646686E-01

FINISHED ITERATION # 1 CHI SQ = 1.2879837E+02

FINISHED ITERATION # 2 CHI SQ = 1.2755936E+02

FINISHED ITERATION # 3 CHI SQ = 1.2667750E+02

FINISHED ITERATION # 4 CHI SQ = 1.2597696E+02

FINISHED ITERATION # 5 CHI SQ = 1.2537410E+02

THERE WERE 5 ITERATIONS

USING $Y=A4+A1/((X-A2)**2 + 0.25*A3 **2)$

YOU OBTAIN:

A(1) = 1.6568207E-03+/- -4.1651741E-06

A(2) = 1.1294275E+03+/- 3.5054553E-03

A(3) = 9.0509854E-03+/- 1.4426374E-05

A(4) = -3.5671993E-01+/- 7.9056752E-02

CHI SQUARED = 1.2537410E+02

TOTAL PRESSURE IS = 53.59 TORR

HWHM = 0.004525 CM-1 OR 135.6709 MHZ

DO YOU WANT A DATA RECAP? YES=1,NO=0?

0

DO YOU WANT TO PLOT DATA ON GRAPHICS ?(YES=1,NO=0)

0

STOP

END OF EXECUTION

CPU TIME: 1.46 ELAPSED TIME: 1:3.32

EXIT

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Table 1. Half-widths (HWHM) and strengths vs. N₂ pressure, for O₃ broadened with O₃ pressure = 3.16 torr at 285 K.

SN ₂ Pressure (torr)	Half-widths		Strengths ^{a, b} cm ⁻² torr ⁻¹
	(mKaysers)	(MHz)	
0.0	1.615	48.43	4.602(-5)
10.15	1.929	57.83	6.738(-5)
15.10	2.290	68.66	4.043(-6)
19.95	2.784	83.47	2.385(-6)
30.30	3.819	114.5	1.435(-6)
50.43	4.491	134.6	7.892(-7)
60.48	4.945	148.3	5.945(-7)
90.80	7.600	227.8	3.853(-7)
100.6	6.834	205.0	3.438(-7)

^aintegration over 55 mKaysers

^bNumbers in () are powers of 10

TDL SPECTROSCOPY FACILITY

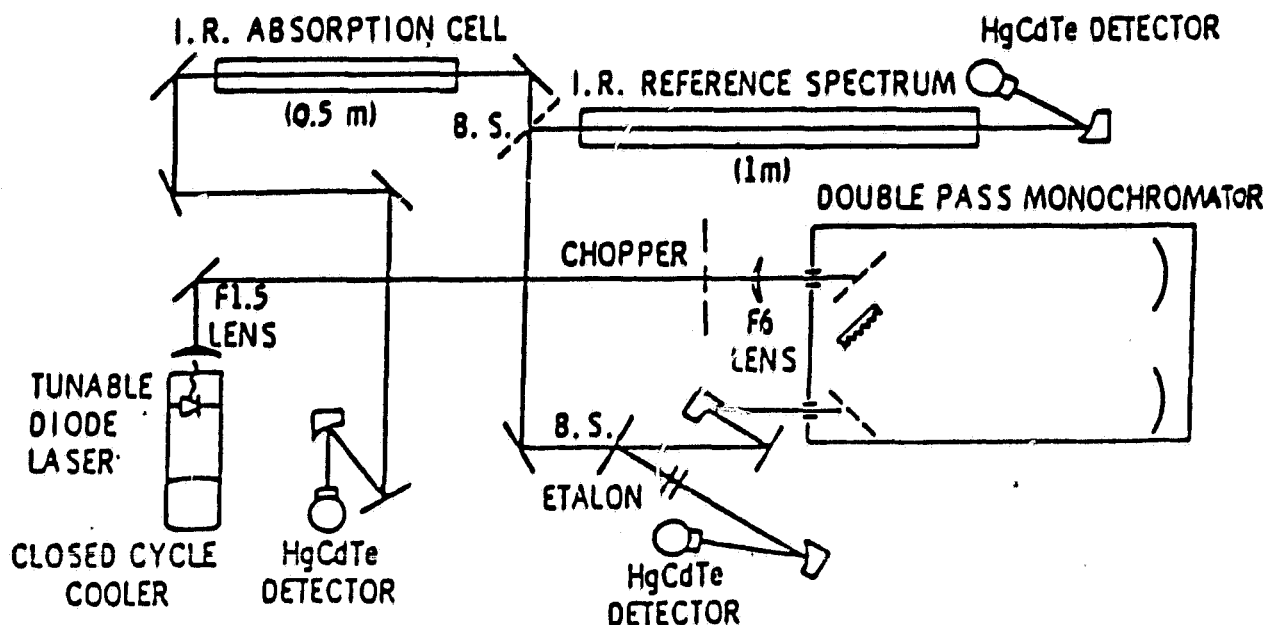


Figure 1. Optical setup for simultaneous measurement of reference gas spectrum, test gas spectrum and etalon tuning curve.

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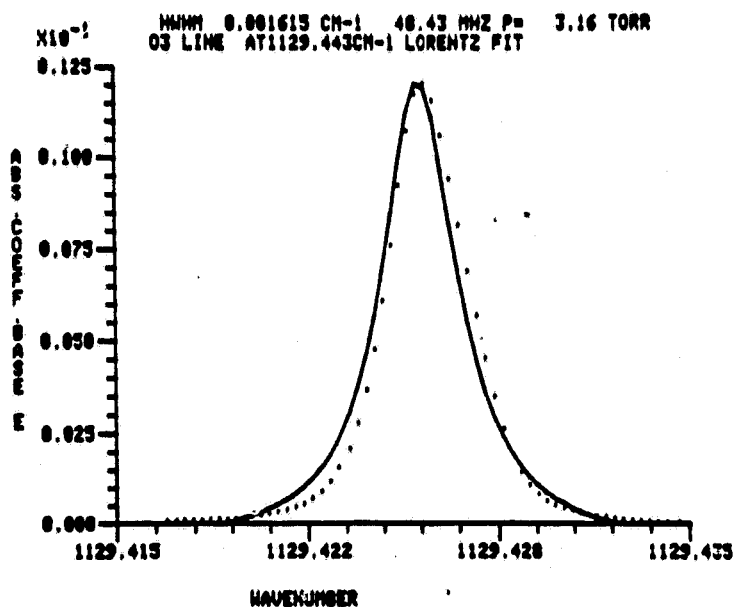


Figure 2. Absorption coefficient vs. wavenumber at pressure = 3.16 torr.

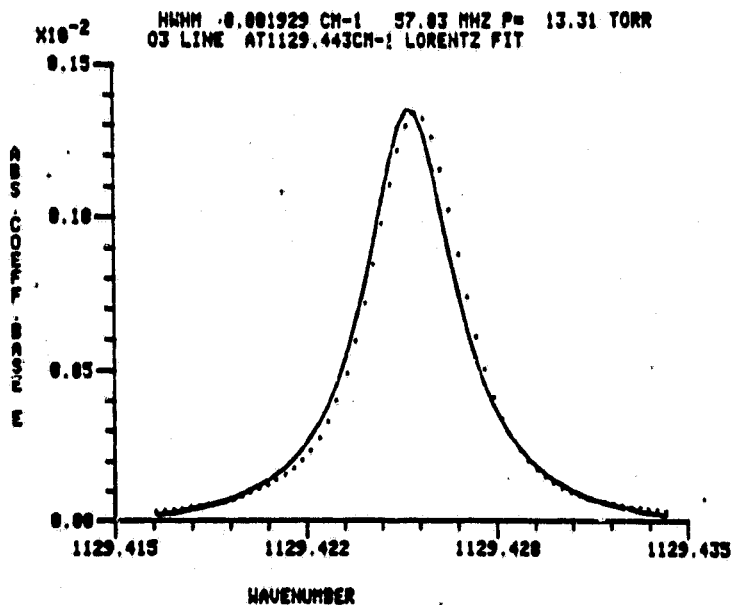


Figure 3. Absorption coefficient vs. wavenumber at pressure = 13.31 torr.

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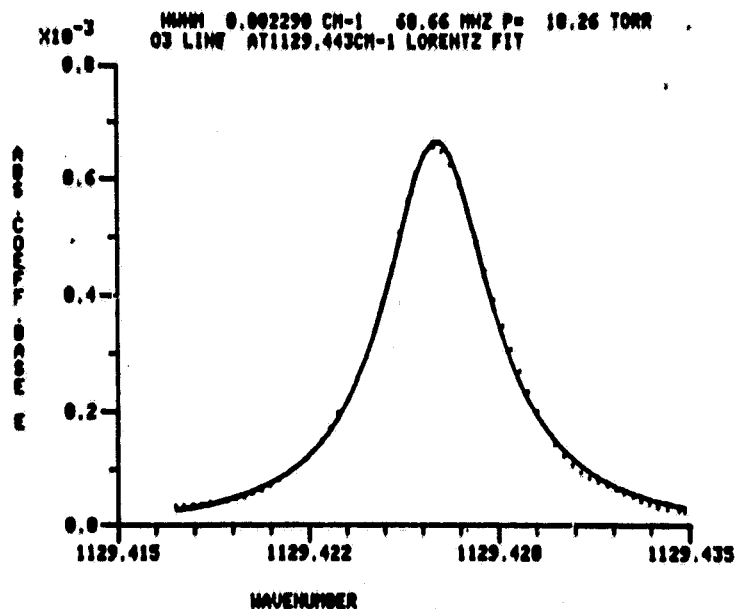


Figure 4. Absorption coefficient vs. wavenumber at pressure = 18.26 torr.

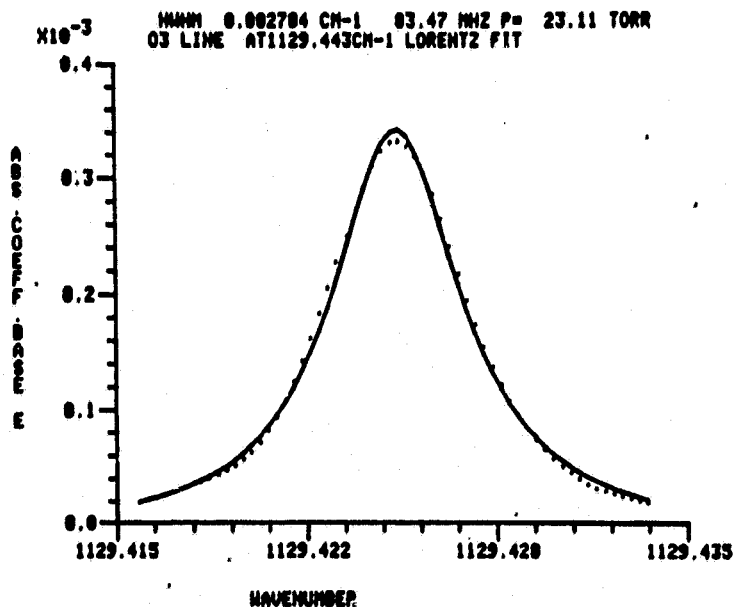


Figure 5. Absorption coefficient vs. wavenumber at pressure = 23.11 torr.

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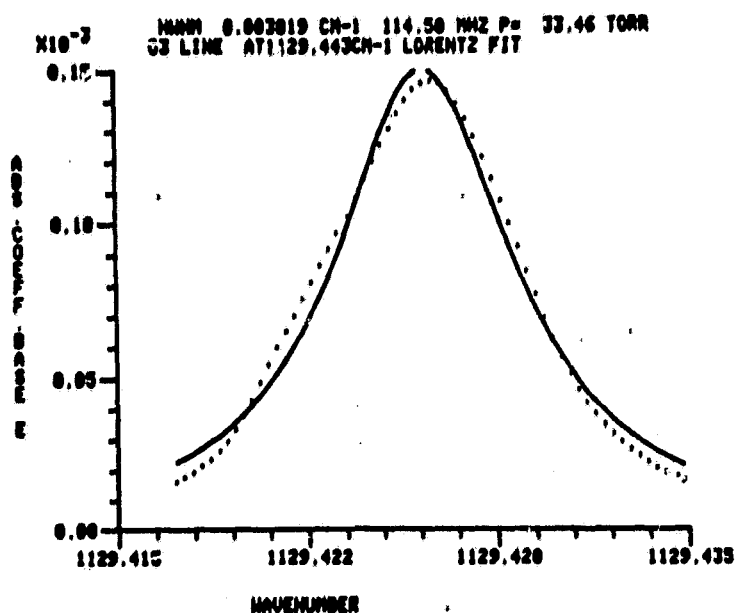


Figure 6. Absorption coefficient vs. wavenumber at pressure = 33.46 torr.

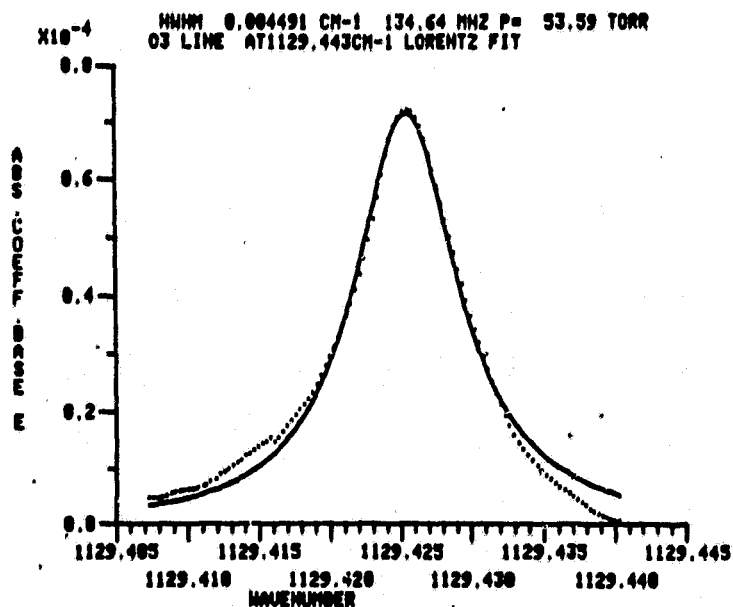


Figure 7. Absorption coefficient vs. wavenumber at pressure = 53.59 torr.

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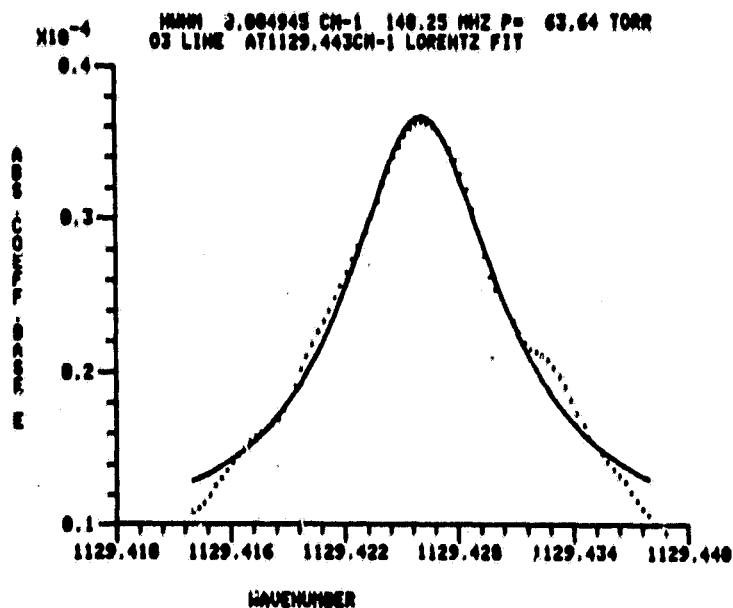


Figure 8. Absorption coefficient vs. wavenumber at pressure = 63.64 torr.

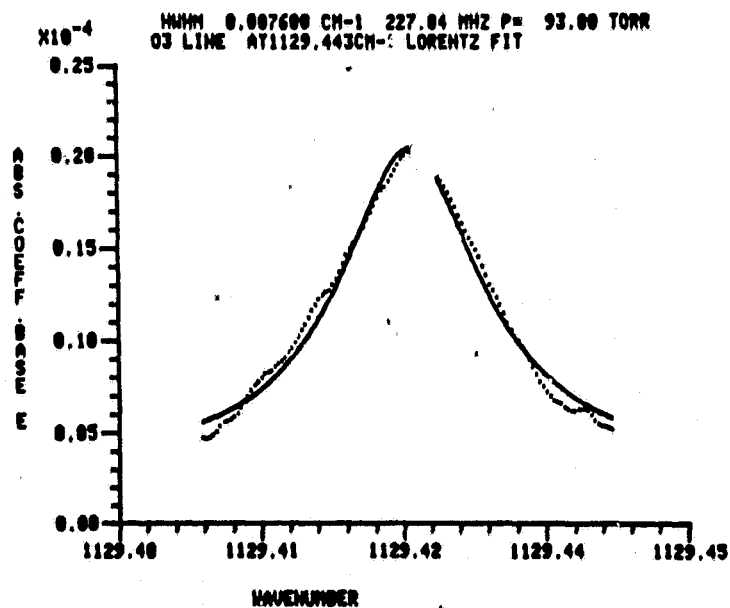


Figure 9. Absorption coefficient vs. wavenumber at pressure = 93.80 torr.

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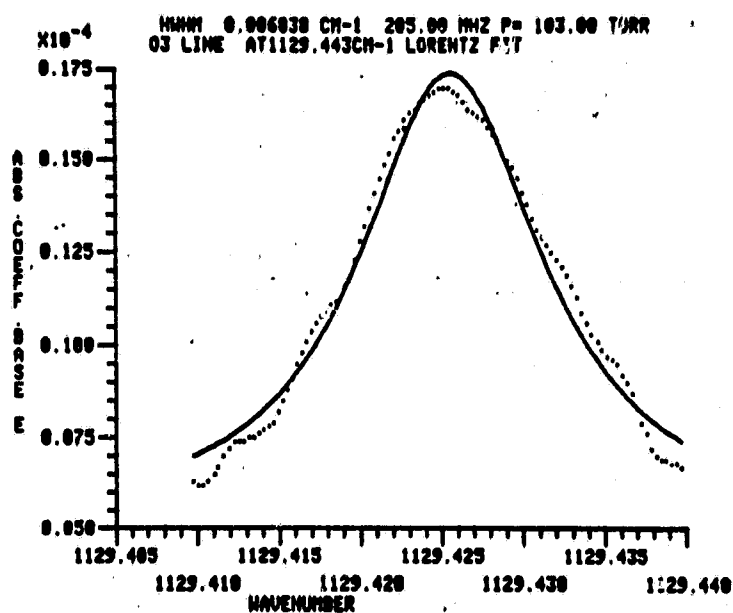


Figure 10. Absorption coefficient vs. wavenumber at pressure = 103.80 torr.

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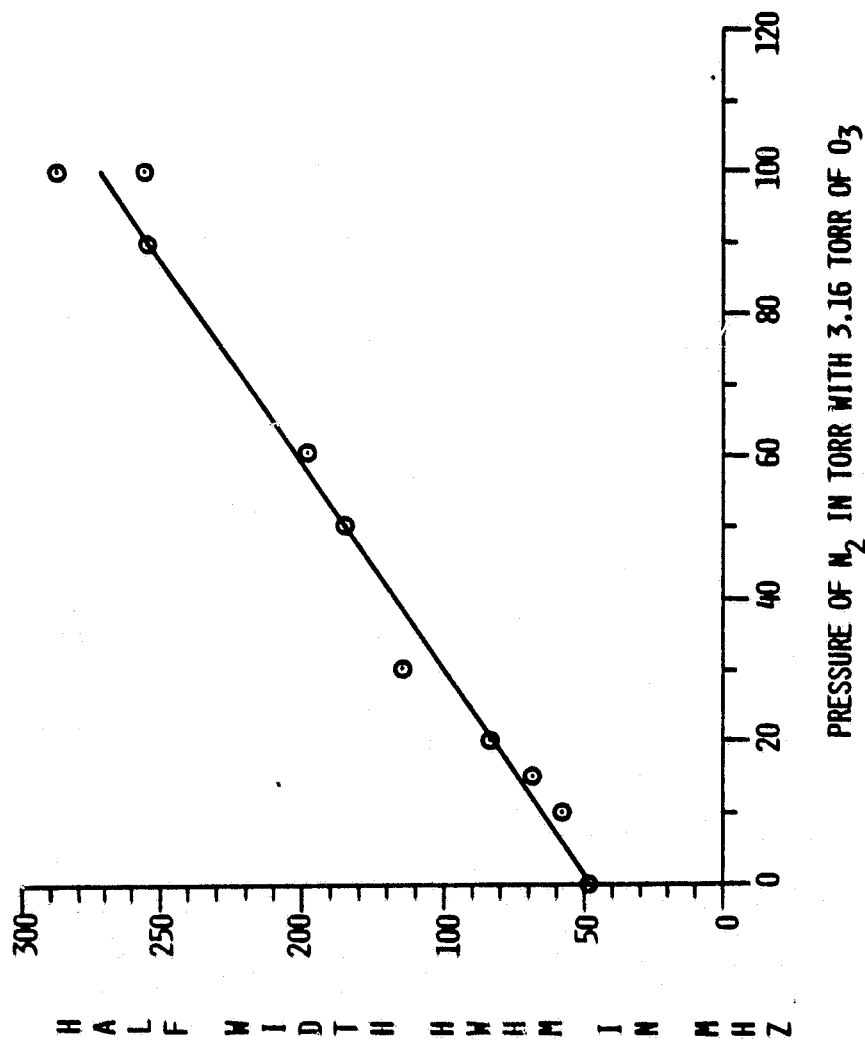


Figure 11. Half-width vs. pressure of N₂ with 3.16 torr of O₃.

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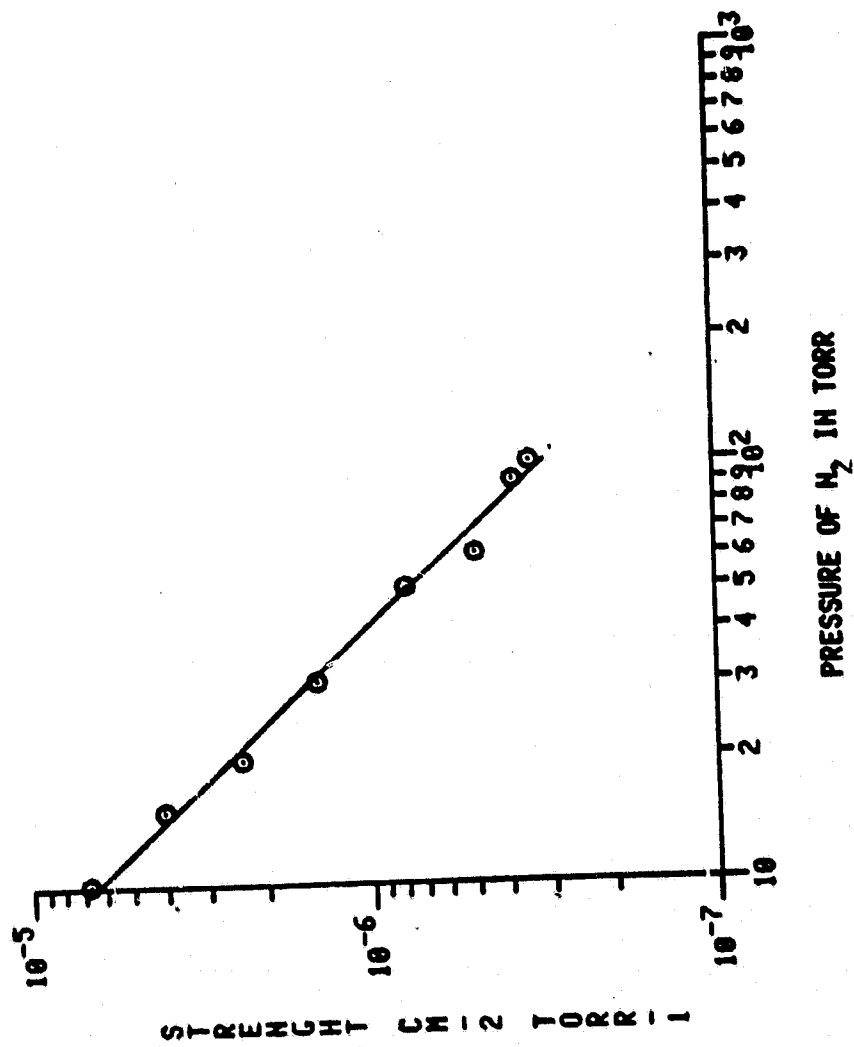


Figure 12. Strength vs. pressure of N₂ for O₃ = 3.16 torr.